

VLBI2010 Simulations at IGG Vienna

Joerg Wresnik, Johannes Boehm,
Andrea Pany, Harald Schuh

Institute of Geodesy and Geophysics (IGG), TU Vienna, Austria

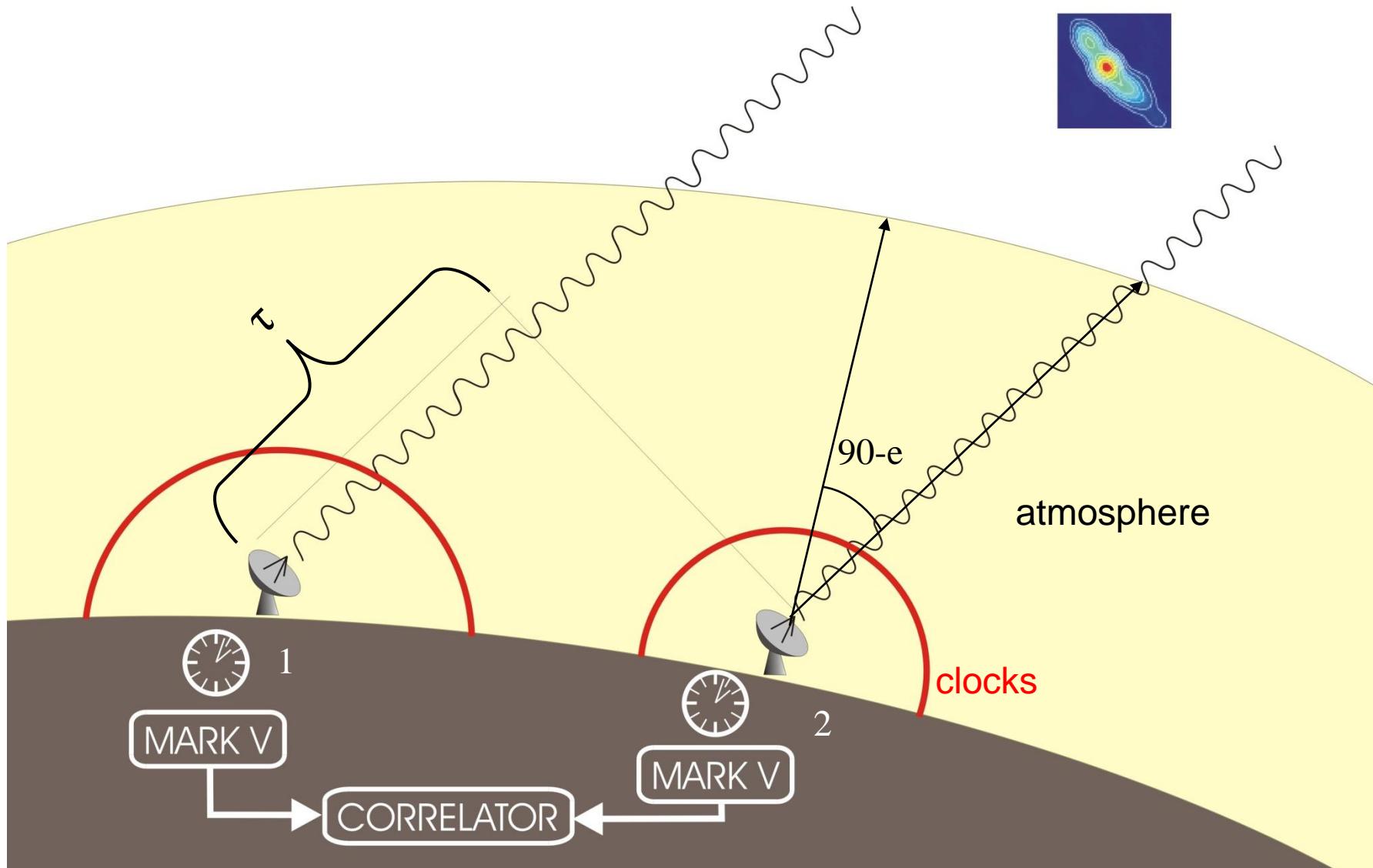
19th European VLBI for Geodesy and
Astrometry (EVGA) Working Meeting,
March, 24 - 25,
Bordeaux, France

- Monte Carlo Simulator
- Evaluation of the Simulator with CONT05
- Network studies
- Slew speed studies for antenna specifications
- Uniform sky scheduling
- Conclusions

Monte Carlo simulation

zwd & clocks are stochastic processes

simulate for station 1 and 2



$$o - c = (zwd_2 \cdot mfw_2(e) + cl_2) - (zwd_1 \cdot mfw_1(e) + cl_1) + wn_{bsl}$$

Simulate the equivalent
wet zenith delay

turbulence model
(by Tobias Nilsson, OSO)

fast turbulence model
(Vienna)

Simulation of zwd and clock parameters

$$o - c = (zwd_2 \cdot mfw_2(e) + cl_2) - (zwd_1 \cdot mfw_1(e) + cl_1) + wn_{bsl}$$

Simulate the equivalent
wet zenith delay

turbulence model
(by Tobias Nilsson, OSO)

fast turbulence model
(Vienna)

Simulate the clocks

random walk + integrated random walk
ASD: 2·10-15@15min
1·10-14@50min

Simulation of zwd and clock parameters

$$o - c = (zwd_2 \cdot mfw_2(e) + cl_2) - (zwd_1 \cdot mfw_1(e) + cl_1) + wn_{bsl}$$

Simulate the equivalent
wet zenith delay

turbulence model
(by Tobias Nilsson, OSO)

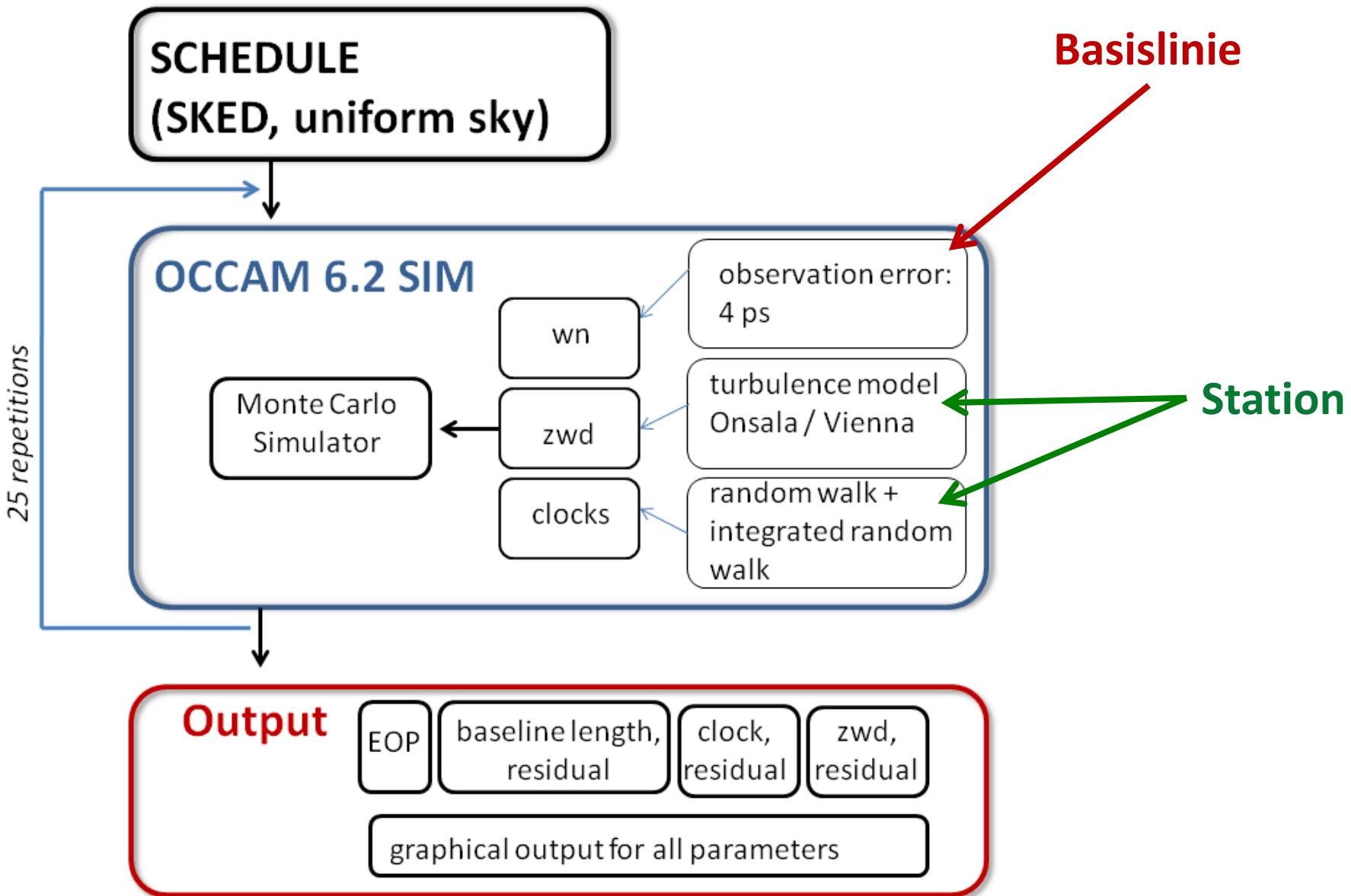
fast turbulence model
(Vienna)

Simulate the clocks

random walk + integrated random walk
ASD: 2·10-15@15min
1·10-14@50min

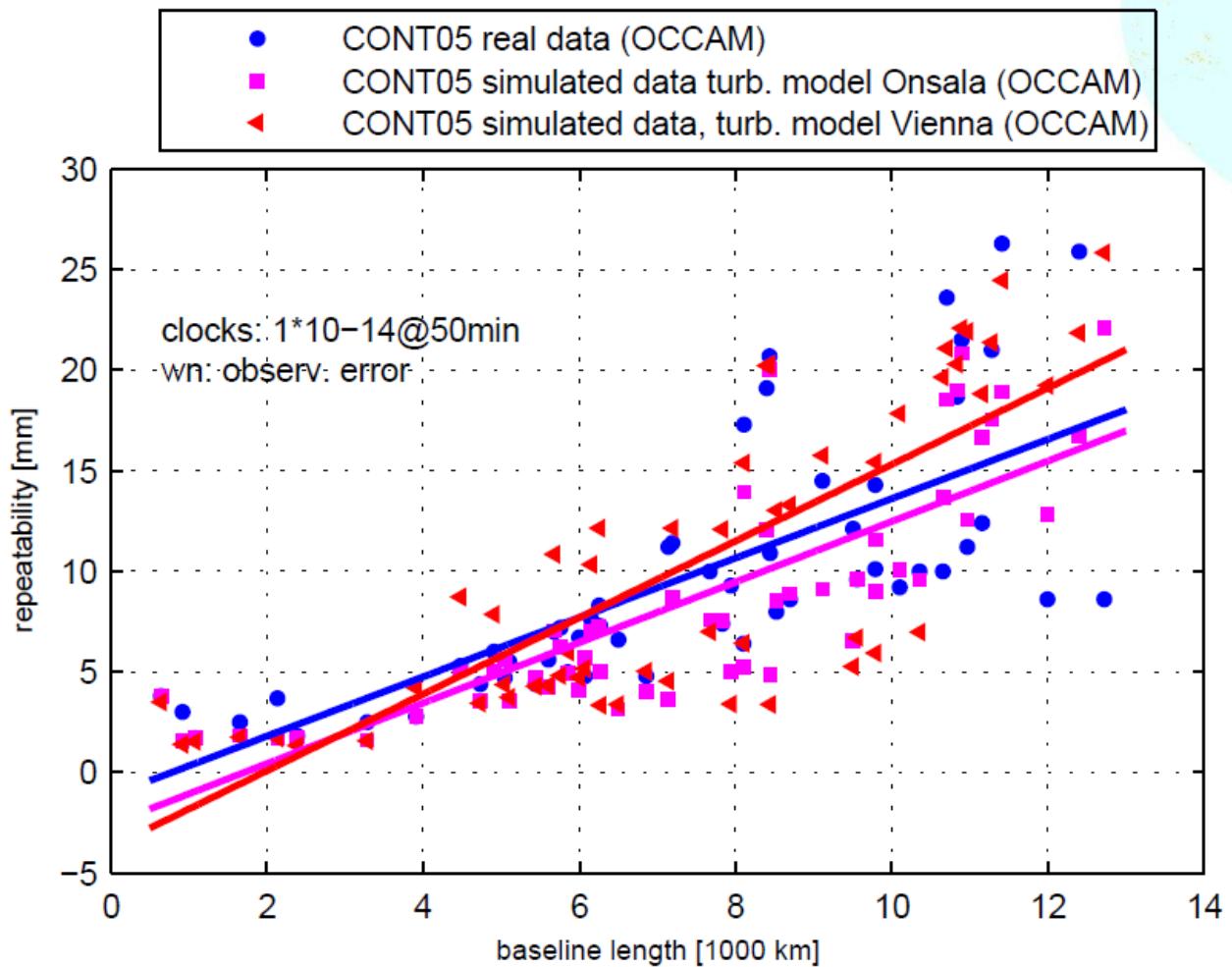
Simulate observation errors

white noise: 4 ps / bsl



Comparison with CONT05

Baseline length repeatability



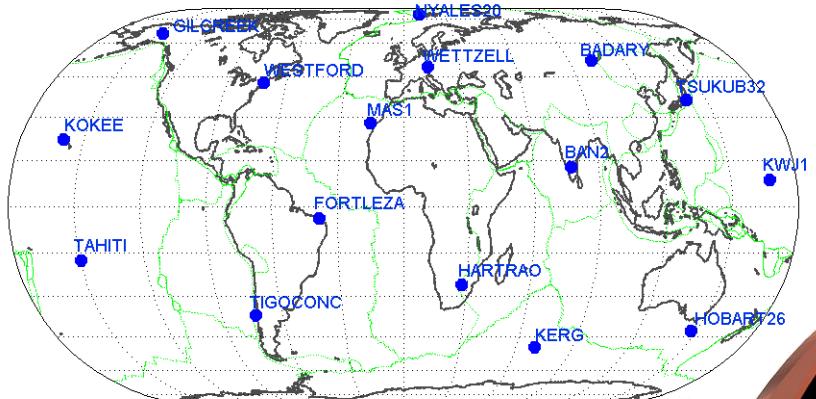
zwd: turbulence model
(Onsala & Vienna)

clocks: $1 \cdot 10^{-14}$ @ 50 min

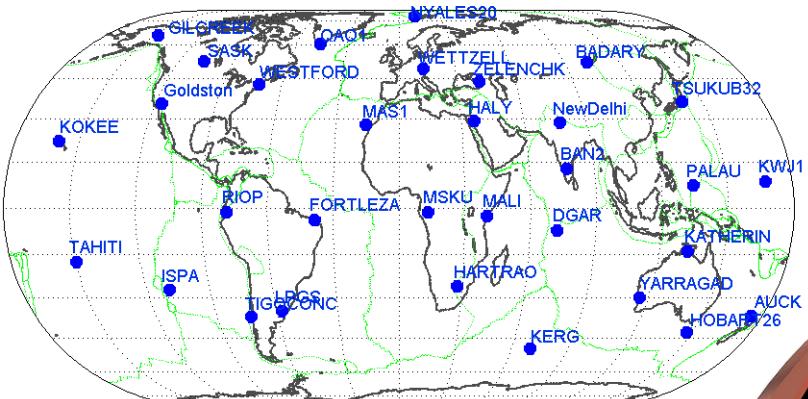
wn: formal errors of
observables

Simulation Networks

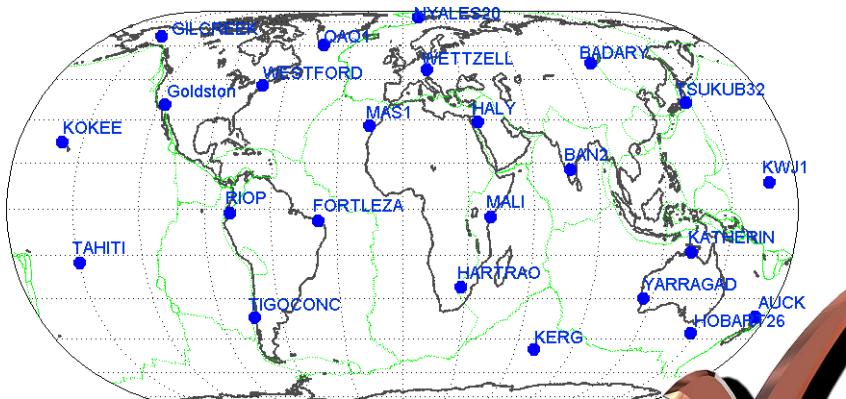
simulation - 16 stns origin(lon,lat) [0 0]



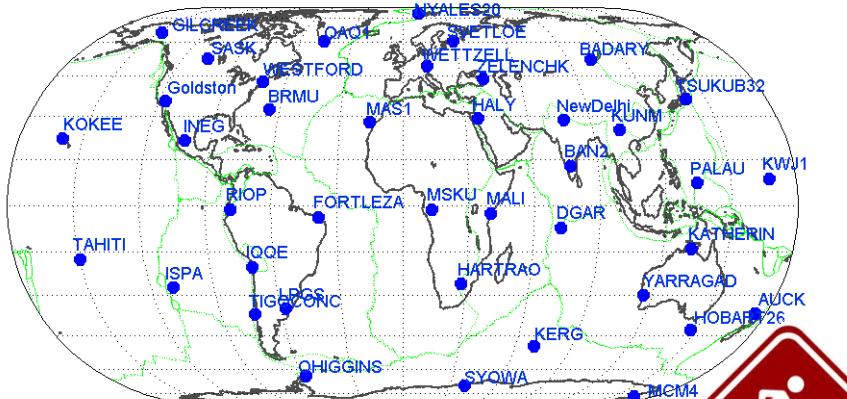
simulation - 32 stns origin(lon,lat) [0 0]



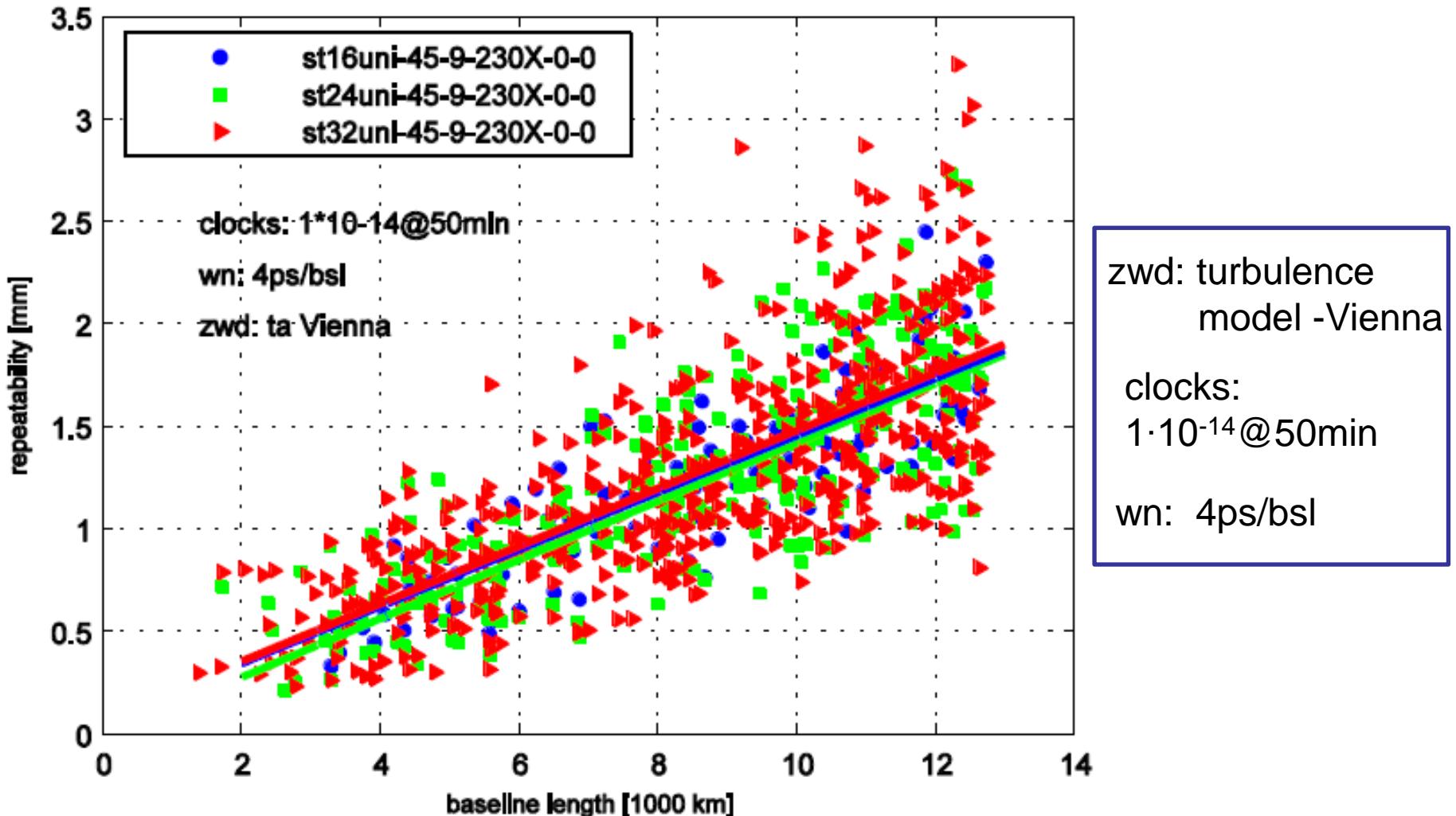
simulation - 24 stns origin(lon,lat) [0 0]



simulation - 40 stns origin(lon,lat) [0 0]

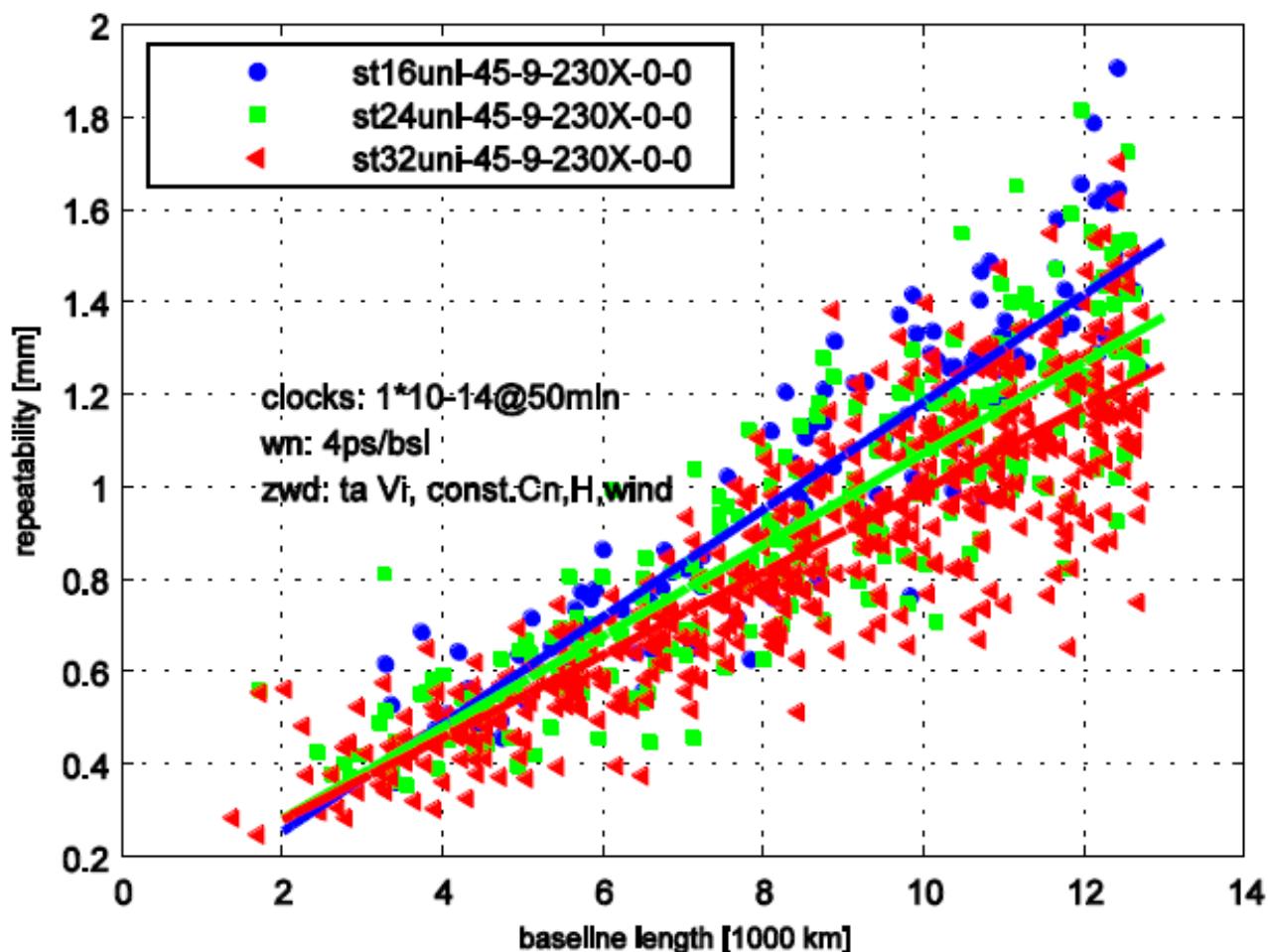


Baseline length repeatability for 16, 24 and 32 station



Baseline length repeatability for 16, 24 and 32 station

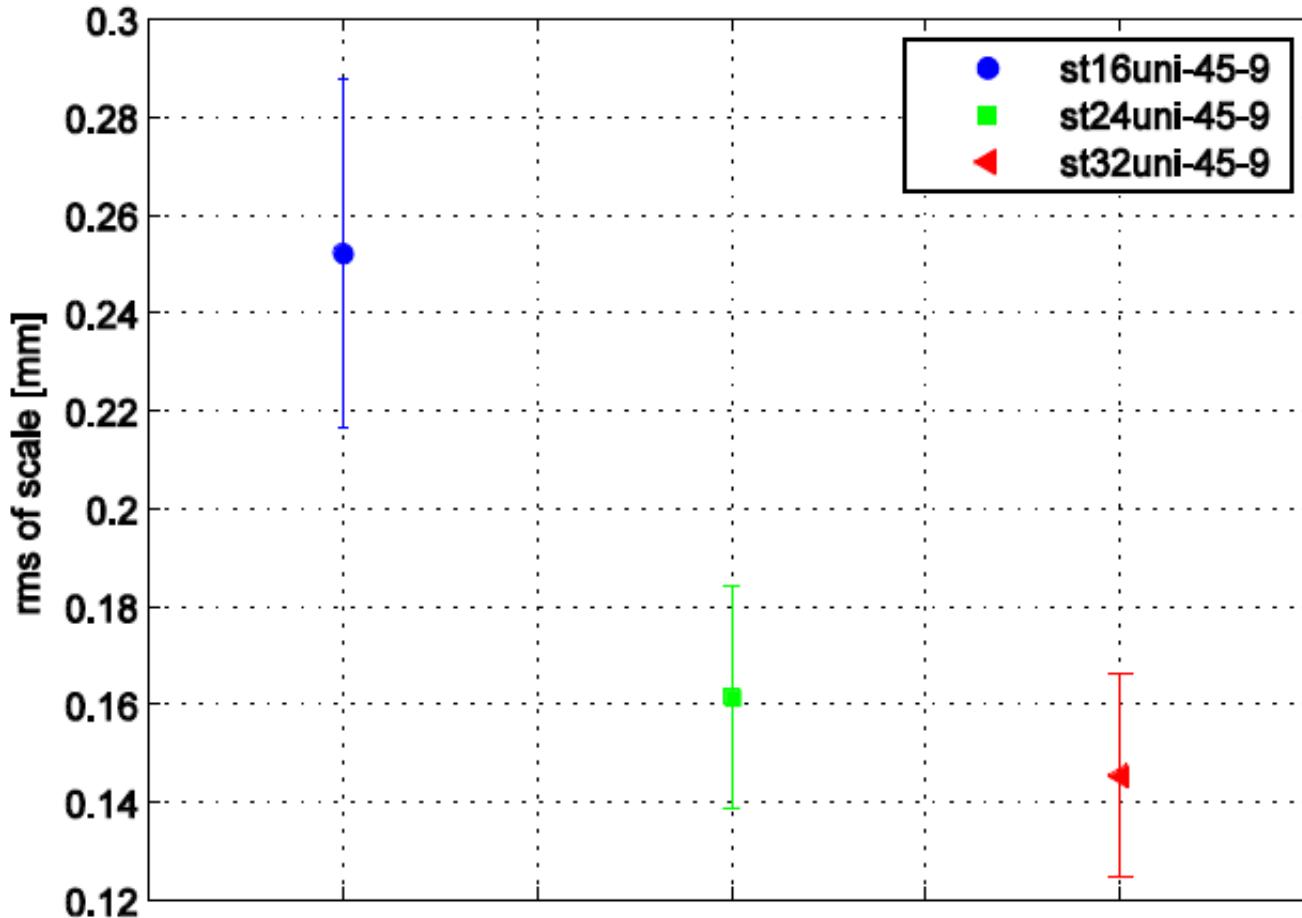
turbulence model uses constant Cn, H and wind speed/direction



zwd: turbulence model –
Vienna
 • $C_n = 2.4 \cdot 10^{-7} \text{ m}^{-1/3}$
 • $H = 1\ 000 \text{ m}$
 • wind = 8 m/s East

clocks:
 $1 \cdot 10^{-14} @ 50\text{min}$
 wn: 4ps/bsl

rms of scale of the 7 parameter Helmert Transformation

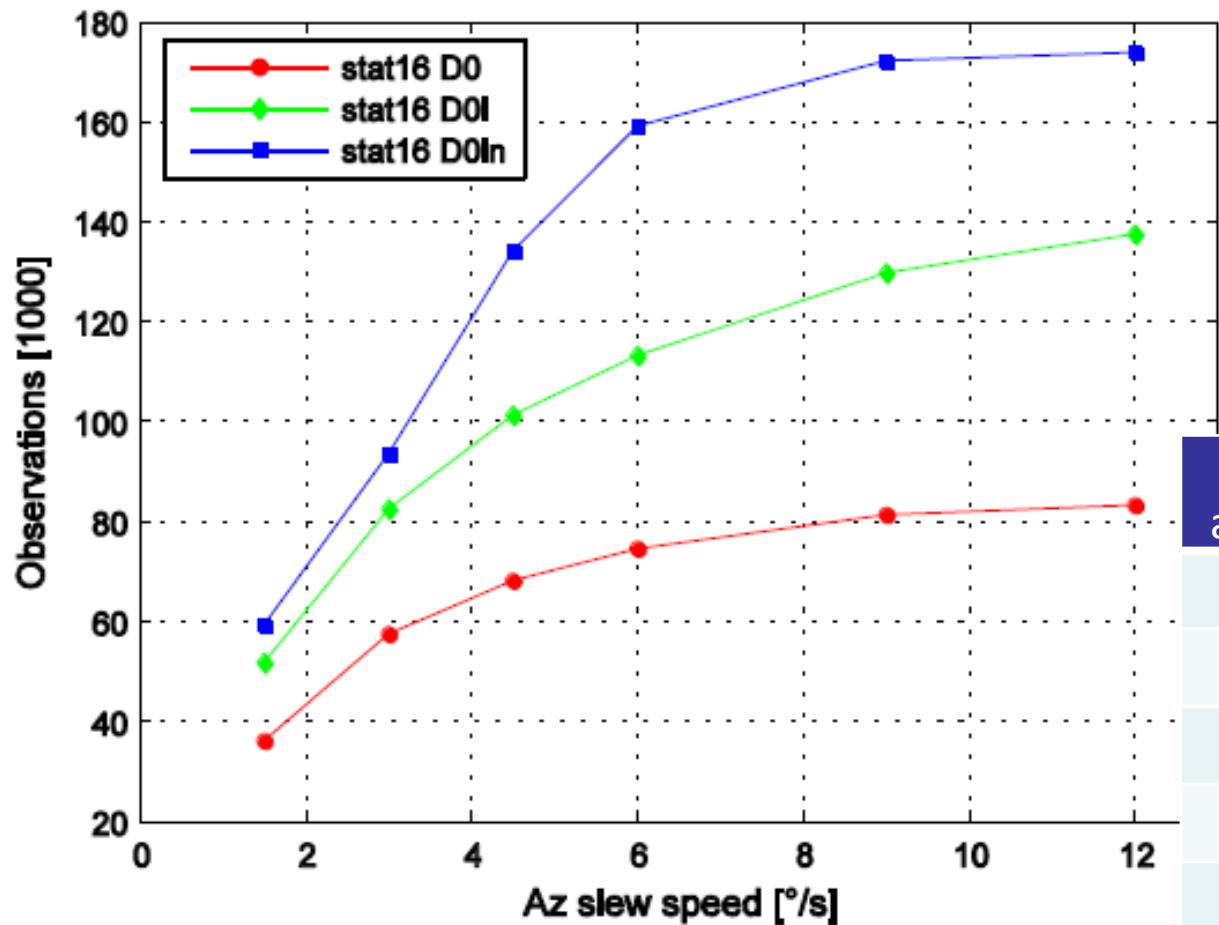


Slew rate tests

D0: geodetic source catalogue, standard optimization

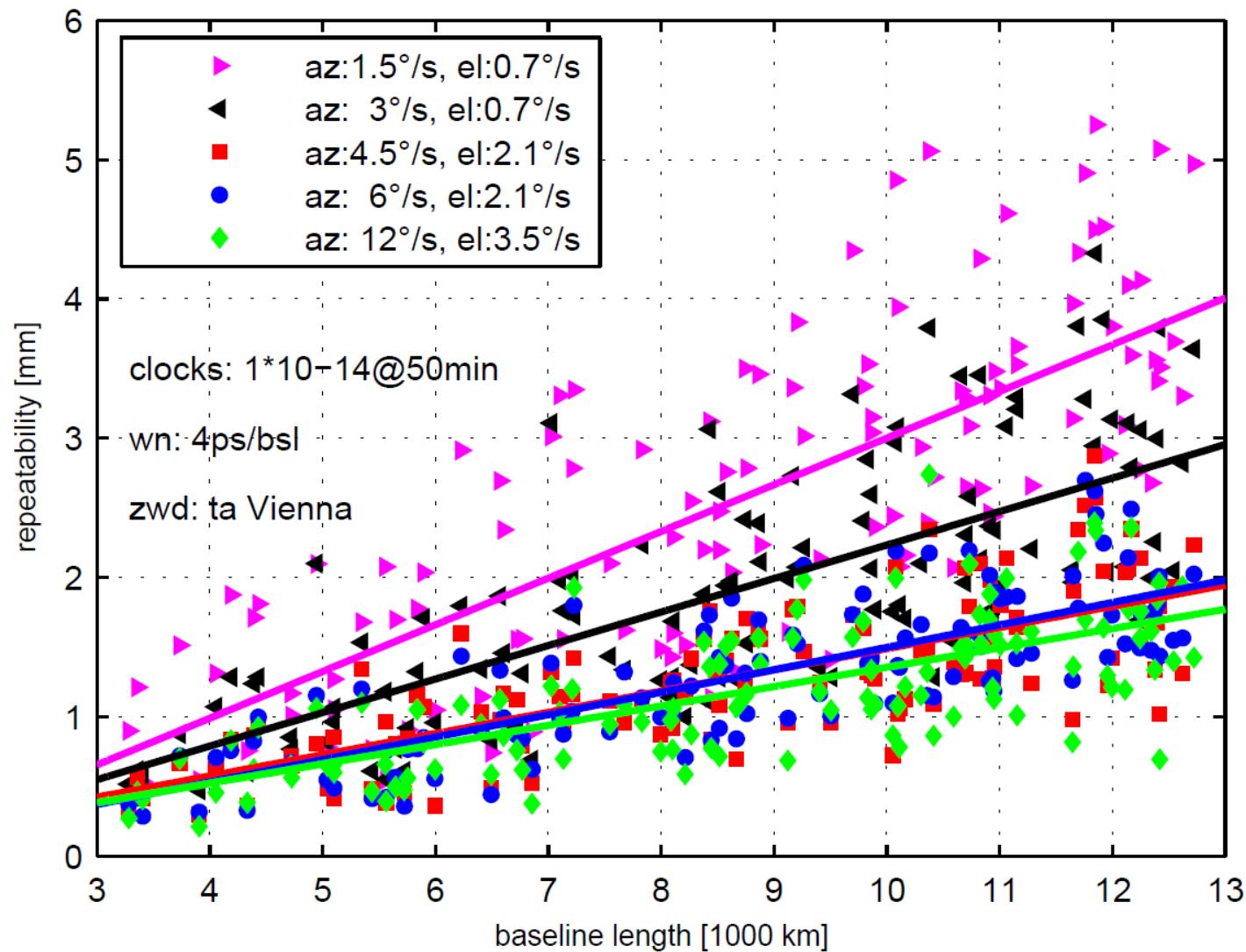
D0I: 230 radio sources, standard optimization

D0In: 230 radio sources, new optimization



slew speed az [°/sec]	el [°/sec]	no. of obs.
12.0	3.5	173831
6.0	2.1	159088
4.5	2.1	134134
3.0	0.7	83149
1.5	0.7	59392

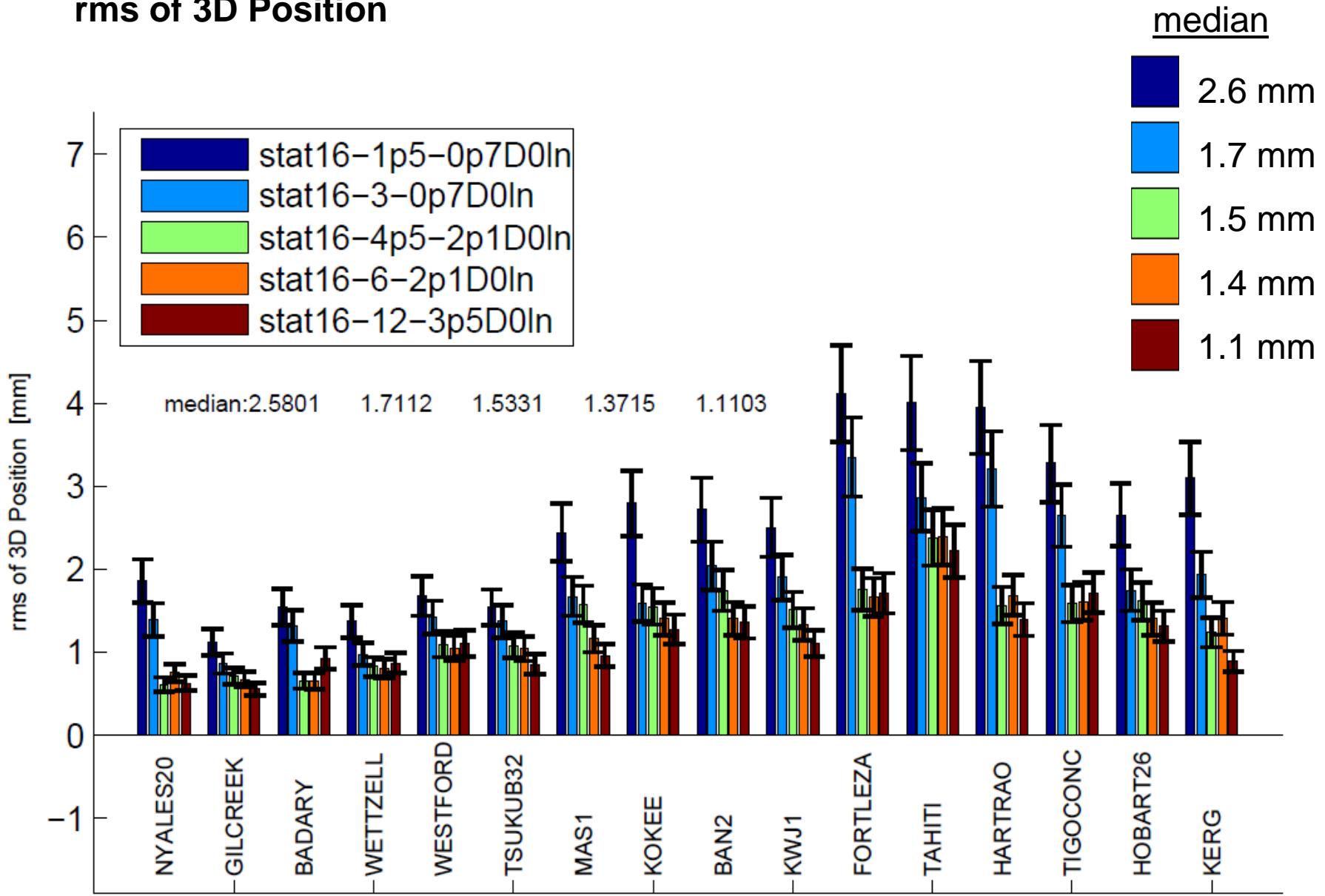
Baseline length repeatability



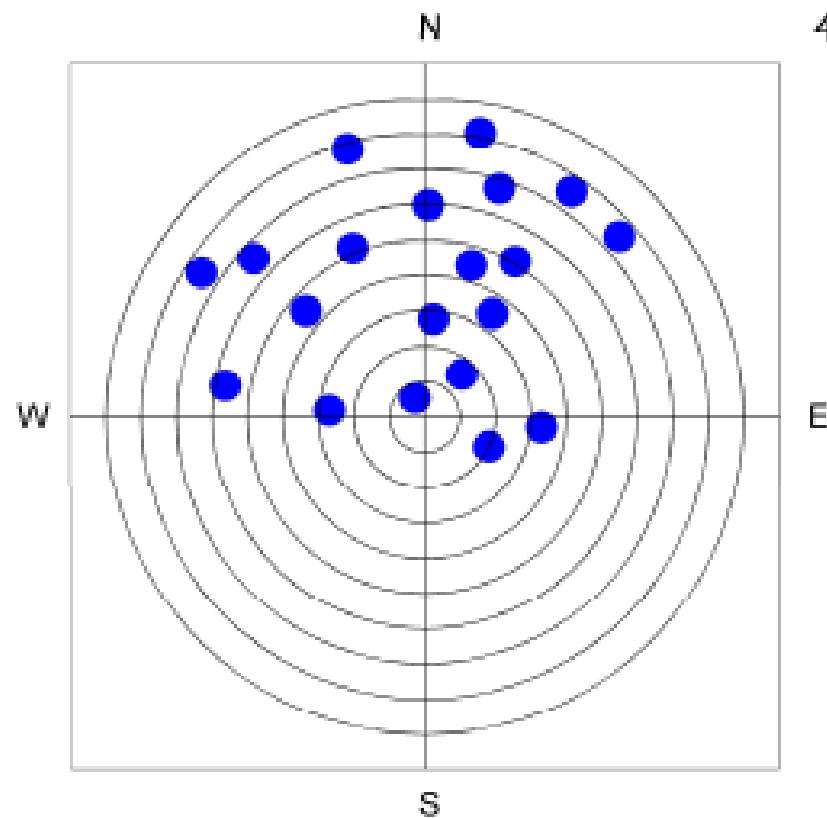
wzd: turbulence model -Vienna
 clocks:
 $1 \cdot 10^{-14}$ @ 50 min
 wn: 4ps/blsl

Slew rate tests

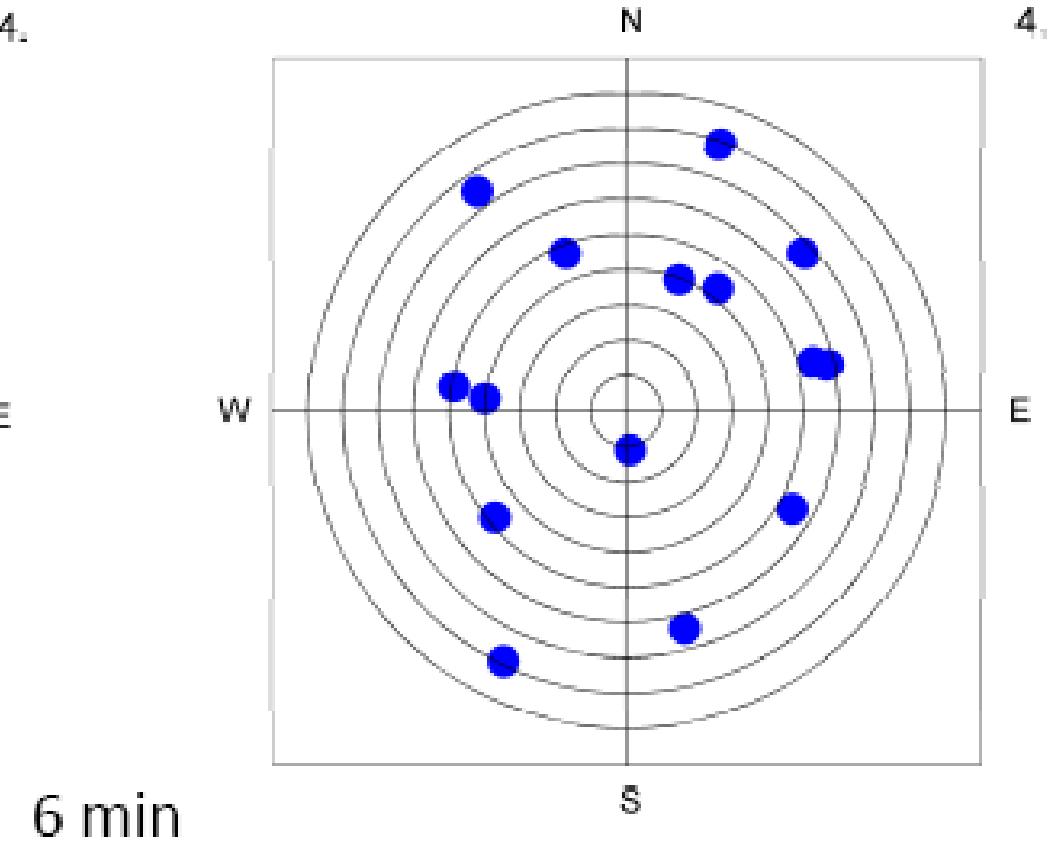
rms of 3D Position



Different scheduling strategies



D0In schedules
observations often clustered



uniform sky schedules
observations well distributed

Uniform sky schedules

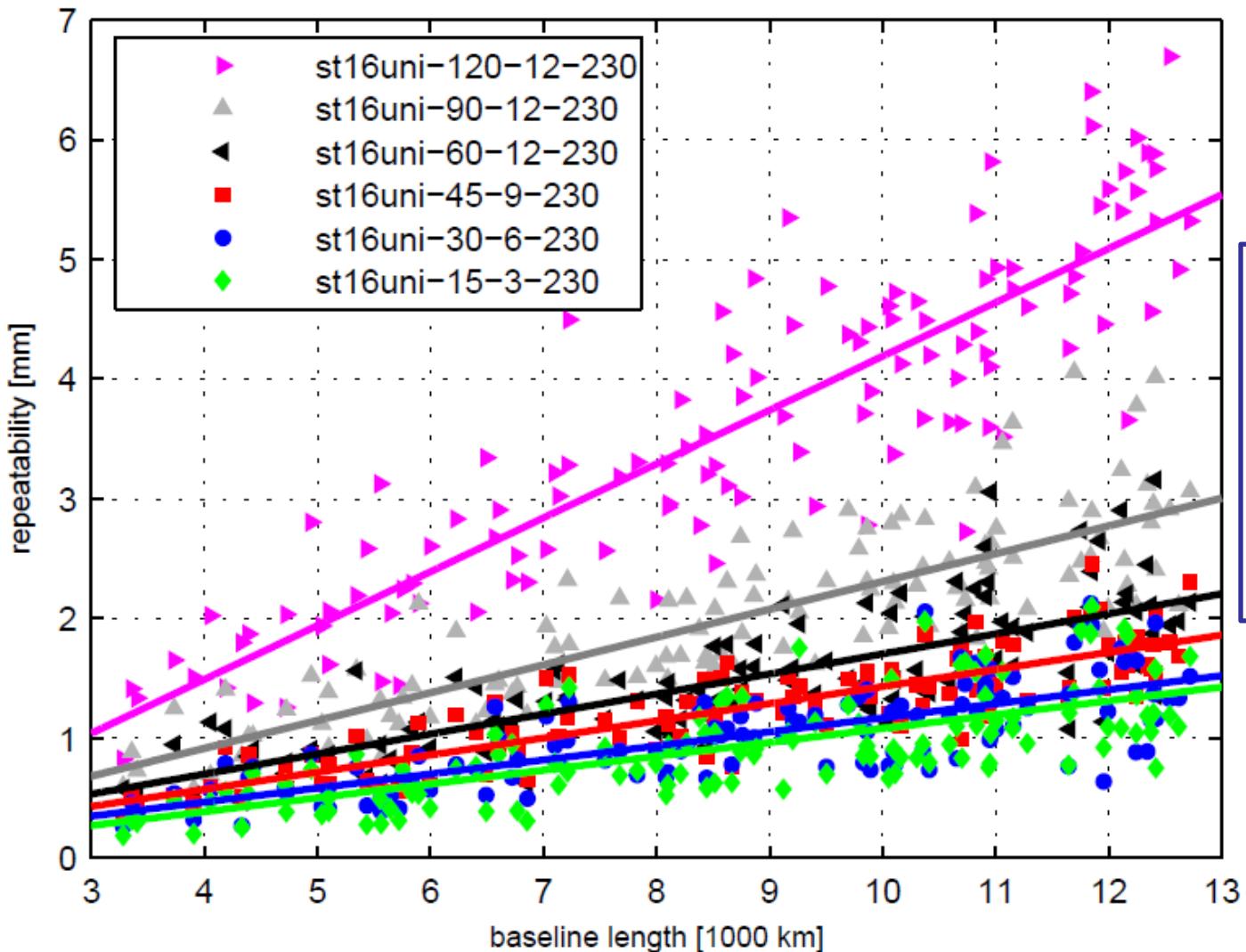
Station	switching intervall	uniform sky	slew speed		no. of obs
			az [°/s]	el [°/s]	
16	15	3	32(**)	8	278 830
16	30	6	12 (*)	3.2	139 564
16	45	9	7.3	1.8	93 231
16	60	12	4.8	1.1	69 708

acceleration: 1.3 °/s^2

* 3 °/s^2

** 8 °/s^2

Baseline length repeatability

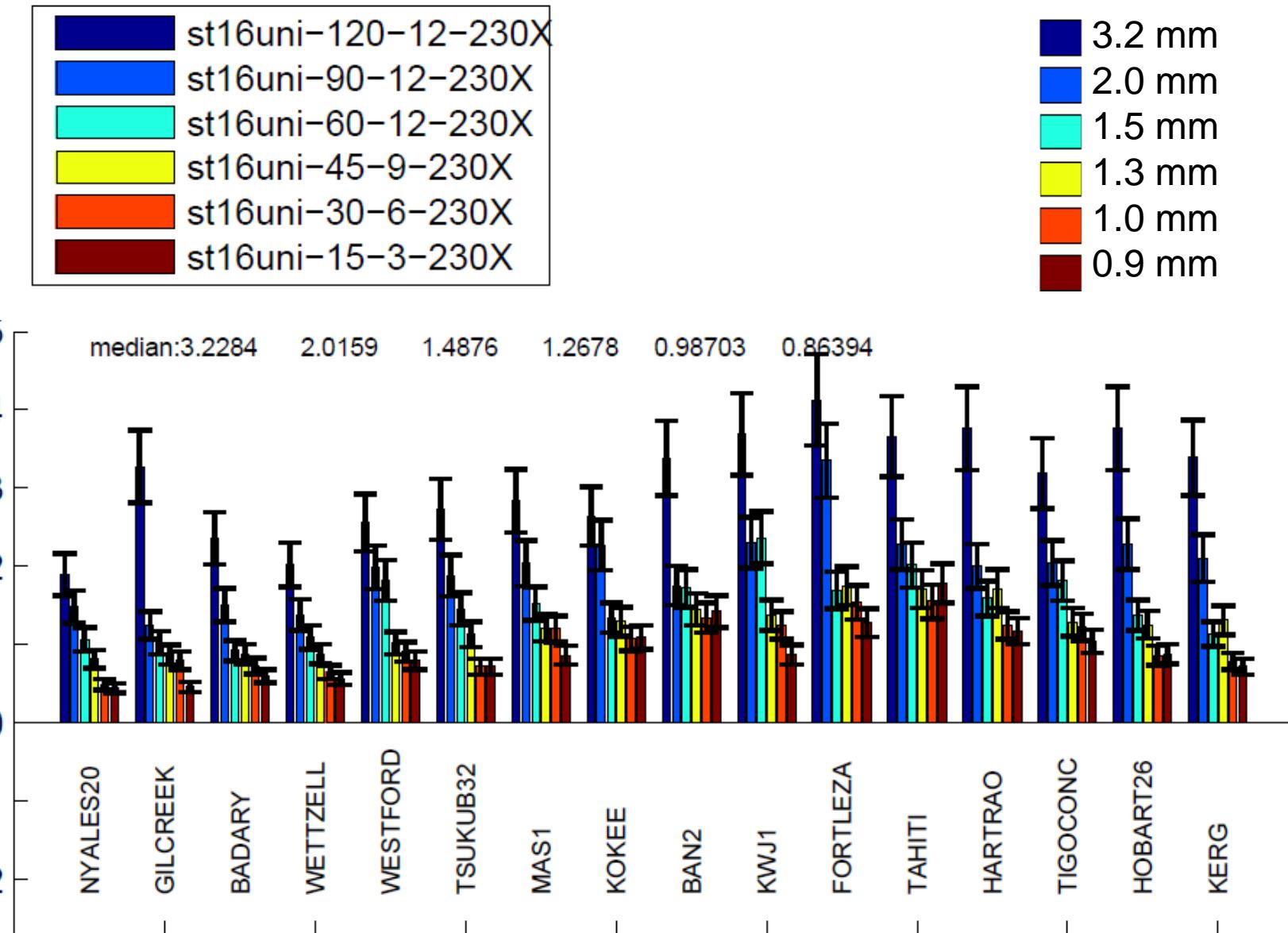


wzd: turbulence model -Vienna
 clocks:
 $1 \cdot 10^{-14} @ 50\text{min}$
 wn: 4ps/bl

Uniform sky schedules

rms of 3D Position

median



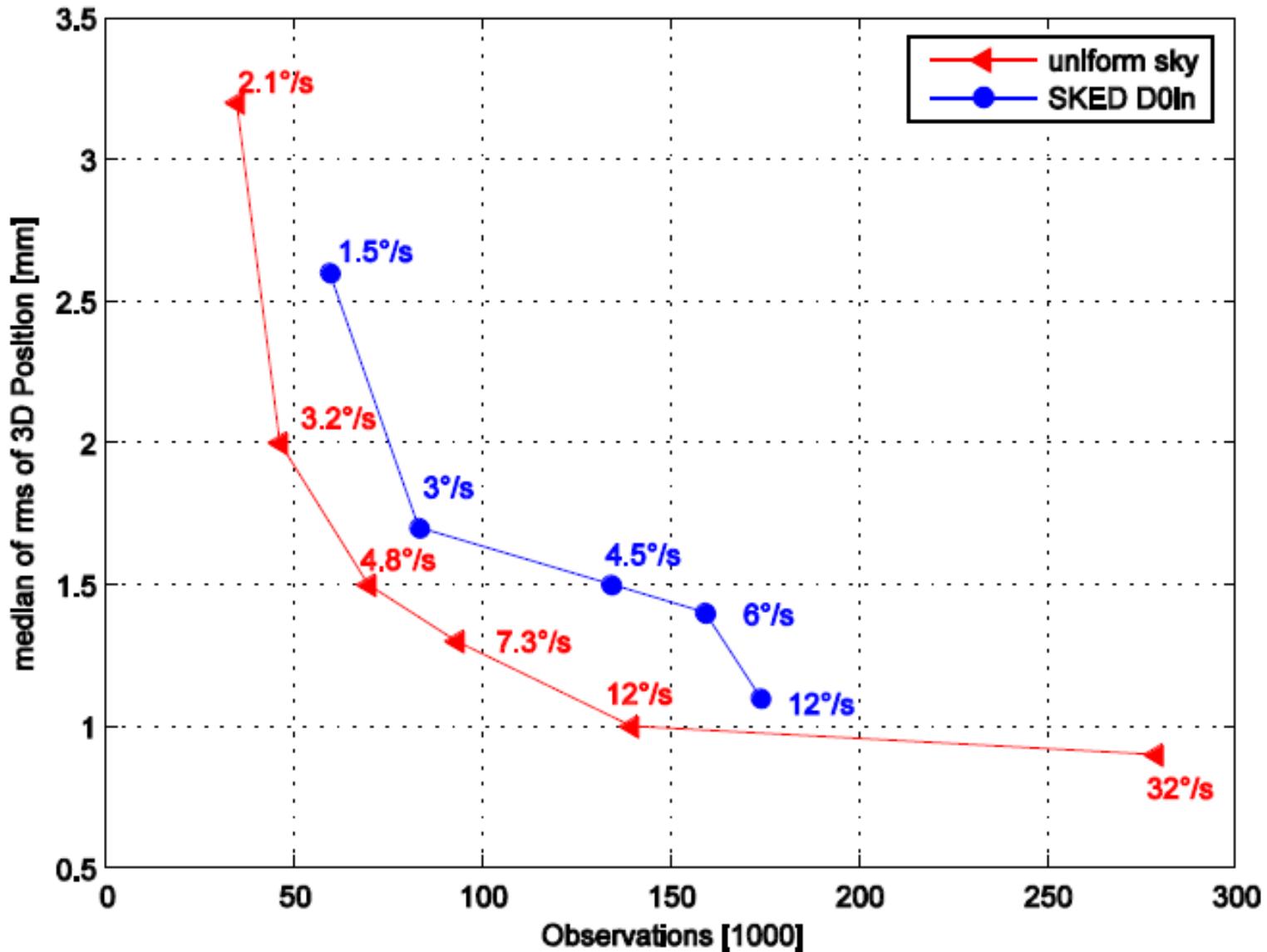
Comparing uniform sky schedule and D0In schedule

Station	switching intervall	uniform sky	slew speed		no. of obs
			az [°/sec]	el [°/sec]	
16	15	3	32(**)	8	278 830
16	30	6	12 (*)	3.2	139 564
16	45	9	7.3	1.8	93 231
16	60	12	4.8	1.1	69 708

D0In

slew speed		no. of obs.
az [°/s]	el [°/s]	
12.0	3.5	173 831
6.0	2.1	159 088
4.5	2.1	134 134
3.0	0.7	83 149
1.5	0.7	59 392

SKED and uniform sky schedules



- The Monte Carlo Simulator shows a good agreement with the CONT05 data.
- The network geometry is important
- The estimation of the zwd is the limiting factor
- The slew rate tests show big improvement for fast antennas $> 6 \text{ } ^\circ/\text{s}$
- The uniform sky scheduling needs fast antennas ($> 6 \text{ } ^\circ/\text{s}$)
- The median of the rms value of 3D position is $< 1 \text{ mm}$ for the uniform sky schedule with the shortest switching interval

Thank you for your
attention!

